

Vladimir S. Matveev (Jena)

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Lorenz Geometry, Granada

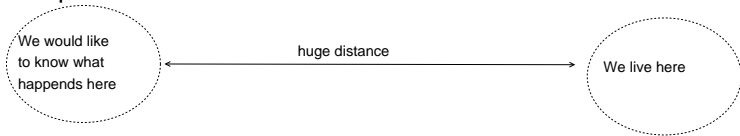
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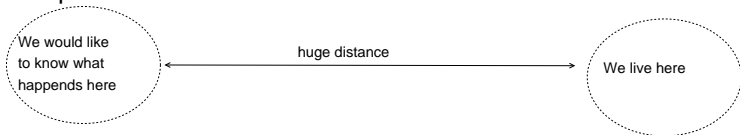
[www.minet.uni-jena.de/~matveev](http://www.minet.uni-jena.de/~matveev)

Suppose we would like to understand the structure of the space-time (i.e., a 4-dimensional metric of Lorenz signature) in a certain part of the universe.



We assume that this part is far enough so the we can use only telescopes (in particular we can not send a space ship there).

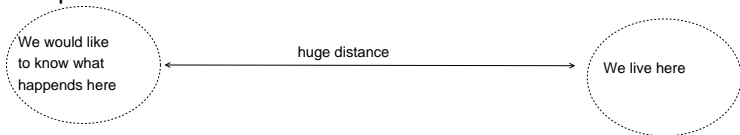
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Then, if the relativistic effects are not negligible (that happens for example if the objects in this part of space-time are sufficiently fast or if this region of the universe is big enough),

**we obtain as a rule the world lines of the objects as unparameterized curves.**

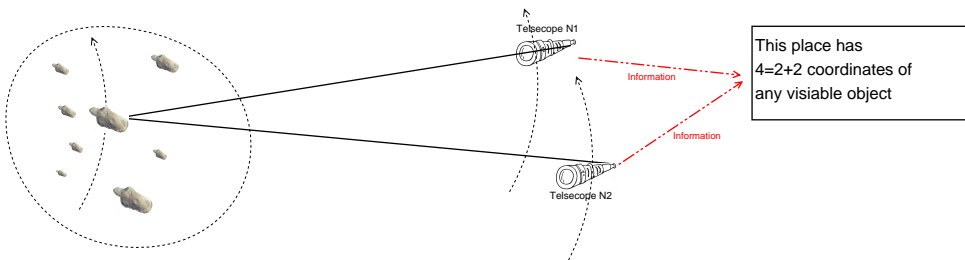
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# In many cases, we can get unparameterized geodesics with the help of astronomic observations

One can obtain unparameterized geodesics by observation:

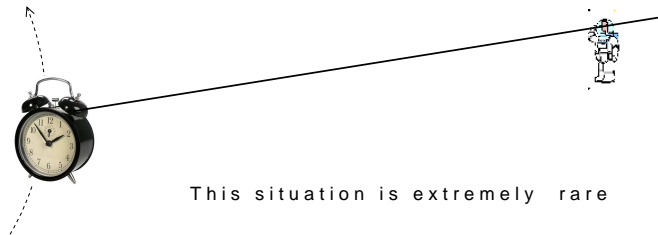


We take 2 freely falling observers that measure two angular coordinates of the visible objects and send this information to one place. This place will have 4 functions  $\text{angle}(t)$  for every visible object which are in the generic case 4 coordinates of the object.

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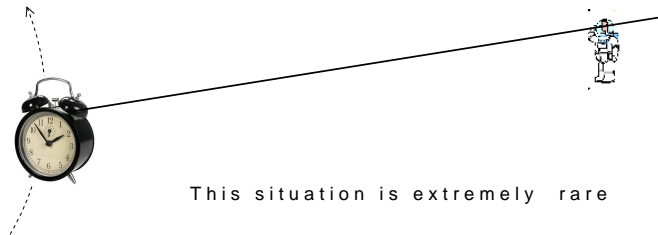
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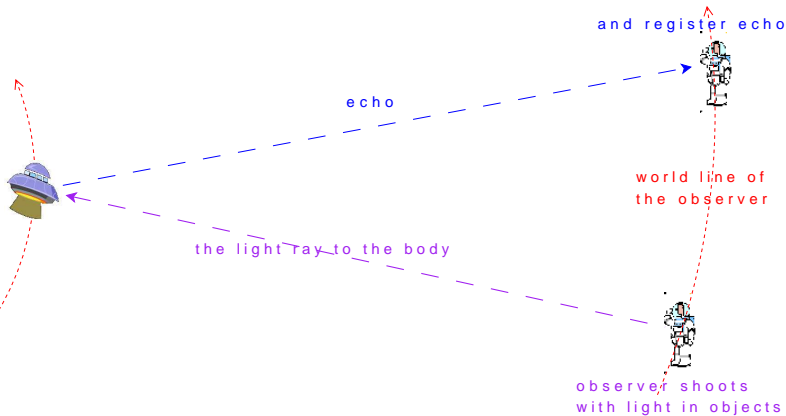
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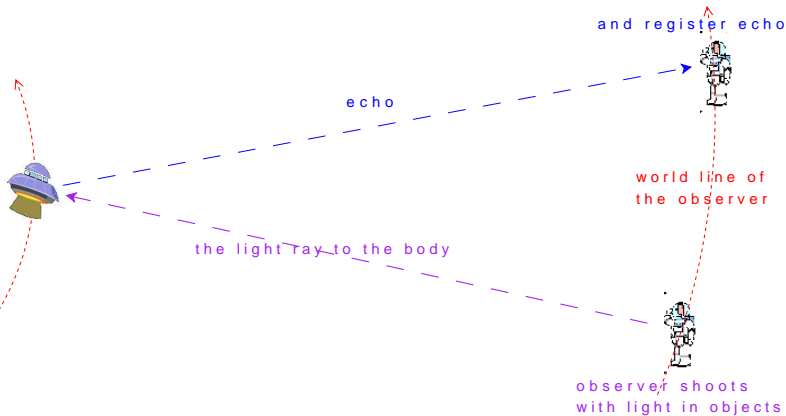
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This approach is related to the so called “**radar coordinates**”. It is a hot topic since 1950th (THE LASER ASTROMETRIC TEST OF RELATIVITY — Space Interferometry Mission), but is applicable in a small neighborhood of solar system only.



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**The mathematical setting:** *We are given a family of smooth curves  $\gamma(t; \alpha)$  in  $U \subseteq \mathbb{R}^4$ ; we assume that the family is sufficiently big in the sense that  $\forall x_0 \in U$*

*$\Omega_{x_0} := \{\xi \in T_{x_0} U \mid \exists \alpha \text{ and } \exists t_0 \text{ with } \frac{d}{dt} \gamma(t; \alpha)|_{t=t_0} \text{ is proportional to } \xi\}$  contains an open subset of  $T_{x_0} U$ .*

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**Jürgen Ehlers 1972**, who said that “We reject clocks as basic tools for setting up the space-time geometry and propose ... freely falling particles instead. We wish to show how the full space-time geometry can be synthesized ... . Not only the measurement of length but also that of time then appears as a derived operation.”



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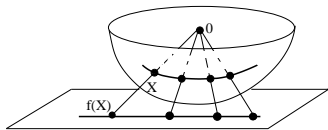
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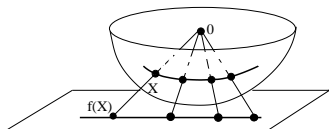
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The example of Lagrange survives for all signatures and for all dimensions.

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**Fact (Dini 1869):** *The metric*

$$(X(x) - Y(y))(dx^2 + dy^2)$$

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**Fact (Levi-Civita 1896)** The metrics of Dini can be generalized for every dimension: The metric

$$-(T(t) - X_1(x_1))(T(t) - X_2(x_2))(T(t) - X_3(x_3))dt^2 + (T(t) - X_1(x_1))(X_1(x_1) - X_2(x_2))(X_1(x_1) - X_3(x_3))dx_1^2 + (T(t) - X_2(x_2))(X_1(x_1) - X_2(x_2))(X_2(x_2) - X_3(x_3))dx_2^2 + (T(t) - X_3(x_3))(X_1(x_1) - X_3(x_3))(X_2(x_2) - X_3(x_3))dx_3^2$$

*is geodesically equivalent to the metric*

$$-\frac{(T(t) - X_1(x_1))(T(t) - X_2(x_2))(T(t) - X_3(x_3))}{T(t)^2 X_1(x_1) X_2(x_2) X_3(x_3)} dt^2 + \frac{(T(t) - X_1(x_1))(X_1(x_1) - X_2(x_2))(X_1(x_1) - X_3(x_3))}{T(t) X_1(x_1)^2 X_2(x_2) X_3(x_3)} dx_1^2 + \frac{(T(t) - X_2(x_2))(X_1(x_1) - X_2(x_2))(X_2(x_2) - X_3(x_3))}{T(t) X_1(x_1) X_2(x_2)^2 X_3(x_3)} dx_2^2 + \frac{(T(t) - X_3(x_3))(X_1(x_1) - X_3(x_3))(X_2(x_2) - X_3(x_3))}{T(t) X_1(x_1) X_2(x_2) X_3(x_3)^2} dx_3^2$$

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## Subproblem 2.1. What metrics 'interesting' for general relativity are geodesically rigid?

**Example – Theorem (Kiosak–Matveev 2009; answers a question explicitly asked by H. Weyl; partial cases are due to Petrov 1961 and Hall-Lonie 2007):** *Let  $(M^4, g)$  be a pseudo-Riemannian Einstein (i.e.,  $\text{Ric} = \frac{\text{Scal}}{4}g$ ) manifold of nonconstant curvature. Then, every  $\bar{g}$  having the same geodesics with  $g$  has the same Levi-Civita connection with  $g$ .*

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**Example.** The so-called Friedman-Lemaitre-Robertson-Walker metric

$$g = -dt^2 + R(t)^2 \frac{dx^2 + dy^2 + dz^2}{1 + \frac{\kappa}{4}(x^2 + y^2 + z^2)} ; \quad \kappa = +1; 0; -1,$$

is not geodesically rigid. Indeed,  $\forall c$  the metric

$$\bar{g} = \frac{-1}{(R(t)^2 + c)^2} dt^2 + \frac{R(t)^2}{c(R(t)^2 + c)} \frac{dx^2 + dy^2 + dz^2}{1 + \frac{\kappa}{4}(x^2 + y^2 + z^2)}$$

is geodesically equivalent to  $g$  (essentially Levi-Civita 1896; repeated by many relativists (Nurowski, Gibbons et al, Hall) later).

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- ▶ In the Riemannian case, such description is due to Levi-Civita 1896.
- ▶ Pseudo-Riemannian case was considered to be solved by Aminova 1993, but recently mathematical difficulties were found in her work.
- ▶ We give an answer (joint with Bolsinov) in dimension 4 and for Lorenz signature of the metric.
- ▶ Actually, we can generalize the answer for all dimensions  $n$  and for all signatures.

# The goal of my talk: to say something about all this (sub)problems

- ▶ **Problem 1. How to reconstruct a metric by its unparameterized geodesics?**
  - ▶ **Subproblem 1.1.** Given a big family of curves  $\gamma(t; a)$ , how to understand whether these curves are reparameterised geodesics of a certain affine connection? How to reconstruct this connection effectively?
  - ▶ **Subproblem 1.2.** Given an affine connection  $\Gamma = \Gamma_{jk}^i$ , how to understand whether there exists a metric  $g$  in the projective class of  $\Gamma$ ? How to reconstruct this metric effectively?
- ▶ **Problem 2.** In what situations is the reconstruction of a metric by the unparameterised geodesics unique (up to the multiplication of the metric by a constant)?
  - ▶ **Subproblem 2.1.** What metrics ‘interesting’ for general relativity are geodesically rigid?
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$\left( \frac{d\gamma(t; \alpha)}{dt} \right)_{|t=t_0}, \left( \frac{d^2 \gamma(t; \alpha)}{dt^2} \right)_{|t=t_0}$ . Since we have infinitely many curves  $\gamma$  passing through  $x_0$ , we have an infinite system of equations.

# Answer to Subproblem 1.1.

**Subproblem 1.** *Given a big family of curves  $\gamma(t; a)$ , how to understand whether these curves are reparameterised geodesics of a certain affine connection? How to construct this connection?*

It is well known (at least since Levi-Civita) that every geodesic  $\gamma : I \rightarrow U$ ,  $\gamma : t \mapsto \gamma^i(t) \in U \subset \mathbb{R}^n$  of  $\Gamma$  is given in terms of **arbitrary parameter**  $t$  as solution of

$$\frac{d^2 \gamma^a}{dt^2} + \Gamma_{bc}^a \frac{d\gamma^b}{dt} \frac{d\gamma^c}{dt} = f \left( \frac{d\gamma}{dt} \right) \frac{d\gamma^a}{dt}. \quad (*)$$

Better known version of this formula assumes that the parameter is affine (we denote it by “ $s$ ”) and reads

$$\frac{d^2 \gamma^a}{ds^2} + \Gamma_{bc}^a \frac{d\gamma^b}{ds} \frac{d\gamma^c}{ds} = 0. \quad (**)$$

Take  $x_0 \in U$ . For every  $\gamma(t; \alpha)$  with  $\gamma(t_0; \alpha) = x_0$  we view the equations  $(*)$  as a system of equations on the entries of  $\Gamma(x_0)$  and on the function  $f|_{\Omega_{x_0}}$ ; the coefficients in this system come from known data

$\left( \frac{d\gamma(t; \alpha)}{dt} \right)_{|t=t_0}, \left( \frac{d^2 \gamma(t; \alpha)}{dt^2} \right)_{|t=t_0}$ . Since we have infinitely many curves  $\gamma$  passing through  $x_0$ , we have an infinite system of equations.

**Theorem (informal version).** *At every point, there exists only one, up to a certain **gauge freedom**, solution  $(\Gamma(x_0)_{jk}^i, f|_{\Omega_{x_0}})$ .*



Gauge freedom (if there exists one solution, there exist many).

$$\text{Repeat: } \frac{d^2\gamma^a}{dt^2} + \Gamma_{bc}^a \frac{d\gamma^b}{dt} \frac{d\gamma^c}{dt} = f \left( \frac{d\gamma}{dt} \right) \frac{d\gamma^a}{dt}. \quad (*)$$

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We consider two connections  $\Gamma$  and  $\bar{\Gamma}$  related by Levi-Civita's formula

$$\Gamma_{bc}^a = \bar{\Gamma}_{bc}^a - \delta_b^a \phi_c - \delta_c^a \phi_b, \quad (1)$$

where  $\phi = \phi_i$  is a one form.

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$$(\delta_b^a \phi_c + \delta_c^a \phi_b) \frac{d\gamma^b}{dt} \frac{d\gamma^c}{dt} = 2 \left( \frac{d\gamma^b}{dt} \phi_b \right) \frac{d\gamma^a}{dt},$$

we obtain that the same curve  $\gamma$  satisfies the equation (17) with respect to the connection  $\bar{\Gamma}$  and the function

$$\bar{f}(v) := f(v) + 2(v^b \phi_b). \quad (2)$$

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Thus, if  $(\Gamma, f)$  is a solution of (\*) (for all  $\gamma$ ), then for every 1-form  $\phi$  the pair  $(\bar{\Gamma}, \bar{f})$  given by (1,2) is also a solution.

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Thus,  $\Gamma$  and  $\bar{\Gamma}$  related by (1) have the same geodesics.

# This is the only gauge freedom

$$\text{Repeat: } \frac{d^2\gamma^a}{dt^2} + \Gamma_{bc}^a \frac{d\gamma^b}{dt} \frac{d\gamma^c}{dt} = f \left( \frac{d\gamma}{dt} \right) \frac{d\gamma^a}{dt}. \quad (*)$$

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We subtract one equation from the other to obtain

$$\tilde{\Gamma}_{bc}^a v^b v^c = -\phi(v) v^a, \quad (3)$$

where  $\tilde{\Gamma} = \bar{\Gamma} - \Gamma$ ,  $\phi = f - \bar{f}$ .



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Since the left hand side is quadratic in  $v$ ,  $\phi$  is linear (i.e.,  $\phi(v) = \phi_i v^i$ ).  
Thus,  $\bar{\Gamma}_{bc}^a - \Gamma_{bc}^a = \delta_b^a \phi_c + \delta_c^a \phi_b$  as we claimed.

# Algorithm how to reconstruct the pair $(\Gamma, f)$ up to the gauge freedom

$$\text{Repeat: } \frac{d^2\gamma^a}{dt^2} + \Gamma_{bc}^a \frac{d\gamma^b}{dt} \frac{d\gamma^c}{dt} = f \left( \frac{d\gamma}{dt} \right) \frac{d\gamma^a}{dt}. \quad (*)$$

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Take a point  $x_0$ ; our goal is to reconstruct  $\Gamma(x_0)_{jk}^i$ . Take  $\gamma(t_0; \alpha)$  such that  $\gamma(t_0; \alpha) = x_0$  and the first component  $\left( \frac{d\gamma^1}{dt} \right)_{|t=t_0} \neq 0$ .

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$$\begin{aligned} f \left( \frac{d\gamma}{dt} \right) &= \left( \frac{d^2\gamma^1}{dt^2} + \Gamma_{ab}^1 \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} \right) / \frac{d\gamma^1}{dt} \\ \frac{d\gamma^2}{dt} \Gamma_{ab}^1 \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} - \frac{d\gamma^1}{dt} \Gamma_{ab}^2 \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} &= \frac{d^2\gamma^2}{dt^2} \frac{d\gamma^1}{dt} - \frac{d\gamma^2}{dt} \frac{d^2\gamma^1}{dt^2} \\ &\vdots \\ \frac{d\gamma^n}{dt} \Gamma_{ab}^1 \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} - \frac{d\gamma^1}{dt} \Gamma_{ab}^n \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} &= \frac{d^2\gamma^n}{dt^2} \frac{d\gamma^1}{dt} - \frac{d\gamma^n}{dt} \frac{d^2\gamma^1}{dt^2}. \end{aligned} \quad (4)$$

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The first equation of (4) is equivalent to the equation  $(*)$  for  $a = 1$  solved with respect to  $f\left(\frac{d\gamma}{dt}\right)$ . We obtain the second, third, etc., equations of (4) by substituting the first equation of (4) in the equations  $(*)$  with  $a = 2, 3$ , etc.

Note that the subsystem of (4) containing the the second, third, etc. equations of (4) does not contain the function  $f$  and is therefore a linear system on  $\Gamma_{jk}^i$ .

$$\begin{aligned}
 \frac{d\gamma^2}{dt} \Gamma_{ab}^1 \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} - \frac{d\gamma^1}{dt} \Gamma_{ab}^2 \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} &= \frac{d^2\gamma^2}{d^2t} \frac{d\gamma^1}{dt} - \frac{d\gamma^2}{dt} \frac{d^2\gamma^1}{d^2t} \\
 &\vdots \\
 \frac{d\gamma^n}{dt} \Gamma_{ab}^1 \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} - \frac{d\gamma^1}{dt} \Gamma_{ab}^n \frac{d\gamma^a}{dt} \frac{d\gamma^b}{dt} &= \frac{d^2\gamma^n}{d^2t} \frac{d\gamma^1}{dt} - \frac{d\gamma^n}{dt} \frac{d^2\gamma^1}{d^2t}.
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Then, for every 'geodesic'  $\gamma(t_0, \alpha)$  gives us  $n - 1$  linear (inhomogeneous) equations on the components  $\Gamma(x_0)_{jk}^i$ .



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Then, for every 'geodesic'  $\gamma(t_0, \alpha)$  gives us  $n - 1$  linear (inhomogeneous) equations on the components  $\Gamma(x_0)_{jk}^i$ . We take a sufficiently big number  $N$  (if  $n = 4$ , it is sufficient to take  $N = 12$ ) and substitute  $N$  generic geodesics  $\gamma(t; \alpha)$  passing through  $x_0$  in this subsystem.

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The system (5) is an overdetermined linear system of PDE of finite type the first order on the unknown functions  $\sigma^{bc}$ ; in theory, there exists an algorithmic method to understand the existence of a solution. The method is highly computational and hardly applicable in this case.

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Weyl has shown that the projective Weyl tensor does not depend of the choice of connection within the projective class: if the connections  $\Gamma$  and  $\bar{\Gamma}$  are related by the formula

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Then, the metric  $\bar{g}$  must satisfy the following system of equations due to the symmetries of the Riemann tensor:

$$\begin{cases} \bar{g}_{ia} W^a_{jkm} + \bar{g}_{ja} W^a_{ikm} = 0 \\ \bar{g}_{ia} W^a_{jkm} - \bar{g}_{ka} W^a_{mij} = 0 \end{cases} \quad (7)$$

The first portion of the equations is due to the symmetry  $(\bar{R}_{ijkm} = -\bar{R}_{jikm})$ , and the second portion is due to the symmetry  $(\bar{R}_{kmij} = \bar{R}_{ijkm})$  of the curvature tensor of  $\bar{g}$ .

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**Theorem (Folklore – Petrov, Hall, Rendall, McIntosh)** Let  $W^i_{jkl}$  be a tensor in  $\mathbb{R}^4$  such that it is skew-symmetric with respect to  $k, \ell$  and such that its traces  $W^a_{akl}$  and  $W^a_{jal}$  vanish. Assume that for all 1-forms  $\xi_i \neq 0$  we have  $W^a_{jkl} \xi_a \neq 0$ . Then, the equations (7) have no more than one-dimensional space of solutions.

Thus, for generic  $\Gamma$ , we can algorithmically reconstruct the conformal class of the metric  $\bar{g}$  by solving the system of linear equations (7). Then, we obtain the ansatz

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## Subproblem 2.1. What metrics are geodesically rigid?

**Theorem (Matveev arXiv:1101.2069)** *Almost every 4D metric is geodesically rigid and can be reconstructed by an algorithm similar to the one above (the algorithm requires solution of linear system of equations and integration).*

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What we understand under almost every? We consider the standard uniform  $C^2$ -topology: the metric  $g$  is  $\varepsilon$ -close to the metric  $\bar{g}$  in this topology, if the components of  $g$  and their first and second derivatives are  $\varepsilon$ -close to that of  $\bar{g}$ .

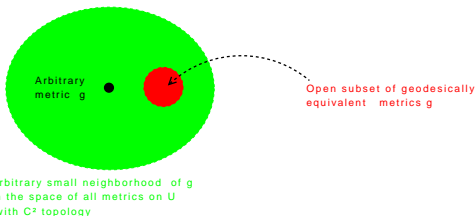


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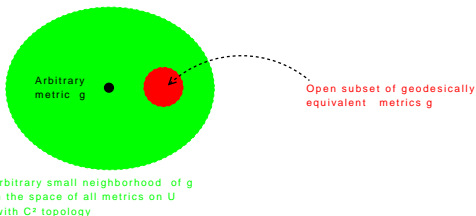


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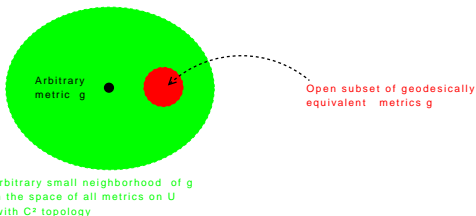
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The result survives for all  $n \geq 4$ . The result survives in D3, if we replace the uniform  $C^2$ -topology by the uniform  $C^3$ -topology. In D2, the result is again true, if we replace the uniform  $C^2$ -topology by the uniform  $C^8$ -topology.

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**Theorem (Gluing Lemma).** Then, the metrics

$$g = \begin{pmatrix} h_1 \chi_2(L_1) & 0 \\ 0 & h_2 \chi_1(L_2) \end{pmatrix}, \quad \bar{g} = \begin{pmatrix} \frac{1}{\chi_2(0)} \bar{h}_1 \chi_2(L_1) & 0 \\ 0 & \frac{1}{\chi_1(0)} \bar{h}_2 \chi_1(L_2) \end{pmatrix}.$$

on  $M_1 \times M_2$ , where  $\chi_i$  is the characteristic polynomial of  $L_i$ , are geodesically equivalent.



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**Example.**

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**Example.** We take two 1-dimensional manifolds

$$\left( I_1, h_1 = dx^2, \bar{h}_1 = \frac{1}{X(x)^2} dx^2 \right) \text{ and } \left( I_2, h_2 = -dy^2, \bar{h}_2 = -\frac{1}{Y(y)^2} dy^2 \right).$$

The corresponding tensors  $L_1$  and  $L_2$  and their characteristic polynomials are

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We see that these metrics are precisely the Dini metrics from the introduction of my talk.

# Building blocks and Splitting Lemma

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The only possible 1-dimensional building block was described above:

$$\left( l_1, h_1 = dx^2, \bar{h}_1 = \frac{1}{X(x)^2} dx^2 \right) \text{ and } \left( l_2, h_2 = \pm dy^2, \bar{h}_2 = \pm \frac{1}{Y(x)^2} dy^2 \right).$$



Two-dimensional building Block were described in Bolsinov-Matveev-Pucacco 2009 (see also Darboux 1886 and Petrov 1949)

	Complex-Liouville case	Jordan-block case
$g$	$\Im(h) dx dy$	$(1 + x Y'(y)) dx dy$
$\bar{g}$	$  \begin{aligned}  & - \left( \frac{\Im(h)}{\Im(h)^2 + \Re(h)^2} \right)^2 dx^2 \\  & + 2 \frac{\Re(h) \Im(h)}{(\Im(h)^2 + \Re(h)^2)^2} dx dy \\  & + \left( \frac{\Im(h)}{\Im(h)^2 + \Re(h)^2} \right)^2 dy^2  \end{aligned}  $	$  \begin{aligned}  & \frac{1 + x Y'(y)}{Y(y)^4} (-2 Y(y) dx dy \\  & + (1 + x Y'(y)) dy^2)  \end{aligned}  $

Trivial block:  $\bar{g} = \text{const} \cdot g$ .

# Three-dimensional building Block were described in Petrov 1949 and Eisenhart 1925

$$\begin{aligned}
 g &= \left( 4 x_2 \left( \frac{d}{dx_3} \lambda(x_3) \right) + 2 \right) dx_1 dx_3 + dx_2^2 \\
 &+ 2 x_1 \left( \frac{d}{dx_3} \lambda(x_3) \right) dx_2 dx_3 + x_1^2 \left( \frac{d}{dx_3} \lambda(x_3) \right)^2 dx_3^2, \\
 \bar{g} &= \frac{1}{\lambda(x_3)^6} \left[ \left( 4 x_2 \lambda(x_3)^2 \left( \frac{d}{dx_3} \lambda(x_3) \right) + 2 \lambda(x_3)^2 \right) dx_1 dx_3 + \lambda(x_3)^2 dx_2^2 \right. \\
 &- \left( 4 x_2 \lambda(x_3) \left( \frac{d}{dx_3} \lambda(x_3) \right) + 2 \lambda(x_3) - 2 x_1 \lambda(x_3)^2 \left( \frac{d}{dx_3} \lambda(x_3) \right) \right) dx_2 dx_3 \\
 &+ \left( 4 x_2^2 \left( \frac{d}{dx_3} \lambda(x_3) \right)^2 + 4 x_2 \left( \frac{d}{dx_3} \lambda(x_3) \right) - 4 x_1 x_2 \lambda(x_3) \left( \frac{d}{dx_3} \lambda(x_3) \right)^2 \right) dx_3^2 \\
 &\left. + \left( 1 + x_1^2 \lambda(x_3)^2 \left( \frac{d}{dx_3} \lambda(x_3) \right)^2 - 2 x_1 \lambda(x_3) \left( \frac{d}{dx_3} \lambda(x_3) \right) \right) dx_3^2 \right]
 \end{aligned}$$

where  $\lambda$  is a function of  $x_3$ , and

$$\begin{aligned}
 g &= 2 dx_3 dx_1 + h(x_2, x_3)_{11} dx_2^2 + 2 h(x_2, x_3)_{12} dx_2 dx_3 + h(x_2, x_3)_{22} dx_3^2, \\
 \bar{g} &= 2 \alpha dx_3 dx_1 + \alpha h(x_2, x_3)_{11} dx_2^2 + 2 \alpha h(x_2, x_3)_{12} dx_2 dx_3 + \beta dx_3^2 + \alpha h(x_2, x_3)_{22} dx_3^2,
 \end{aligned}$$

where  $\alpha$  and  $\beta$  are constants.

# Success report

We have described explicitly all building blocks that can be used in constructing 4D metrics of Lorenz signature; Splitting-Gluing Lemmas give us the explicit construction.

# Success report

We have described explicitly all building blocks that can be used in constructing 4D metrics of Lorenz signature; Splitting-Gluing Lemmas give us the explicit construction. Let us count the number of cases: we can represent 4 as the sum of natural numbers by 4 different ways:

Dim of blocks	Description of blocks	# of cases
1+1+1+1	All building blocks are one-dimensional, and the metric is essentially the Levi-Civita metric from the introduction	1
1+1+2	The first two building blocks are one-dimensional, the third is two-dimensional	3
2+2	Both building blocks are two-dimensional; at least one of them is trivial (i.e., $\bar{h} = \text{const} \cdot h$ )	3
1+3	The first building block is one-dimensional, the second is three-dimensional ('Petrov', 'Eisenhart', or trivial)	3

# Summary: I said something about all this (sub)problems

- ▶ **Problem 1. How to reconstruct a metric by its unparameterized geodesics?**
  - ▶ **Subproblem 1.1.** Given a big family of curves  $\gamma(t; a)$ , how to understand whether these curves are reparameterised geodesics of a certain affine connection? How to reconstruct this connection effectively? **Solved completely**
  - ▶ **Subproblem 1.2.** Given an affine connection  $\Gamma = \Gamma_{jk}^i$ , how to understand whether there exists a metric  $g$  in the projective class of  $\Gamma$ ? How to reconstruct this metric effectively?  
**Suggested an effective way for Ricci-flat metrics**
- ▶ **Problem 2.** In what situations is the reconstruction of a metric by the unparameterised geodesics unique (up to the multiplication of the metric by a constant)?
  - ▶ **Subproblem 2.1.** What metrics 'interesting' for general relativity are geodesically rigid? **Almost every metric is geodesically rigid**
  - ▶ **Subproblem 2.2.** Construct all pairs of nonproportional geodesically equivalent metrics. **Solved completely**

Thank you!!!